

Regional Diamond Exploration Under Cover



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150 years of modern diamond exploration has turned up

7000 kimberlites

1000 diamondiferous

60 economic

10 Tier-1 (\$20b reserve)





Measure potential fields:
EM conductivity
Mag susceptibility
Gravity density

Geochemical samples

Thermal does both and
can penetrate moderate
vegetation and
transported material



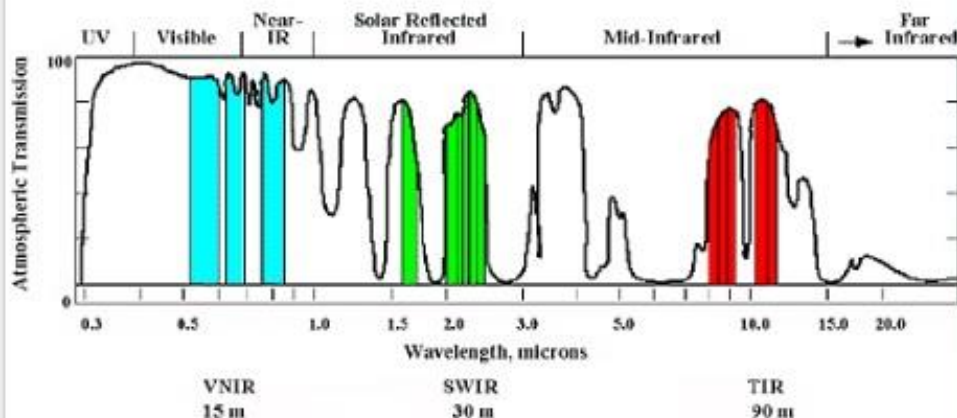
Japanese Aster satellite launched late 1999 images 5 bands thermal @ 90m

Japanese Aster satellite

First images collected in 2000
Still operating



ASTER Spectral Bands



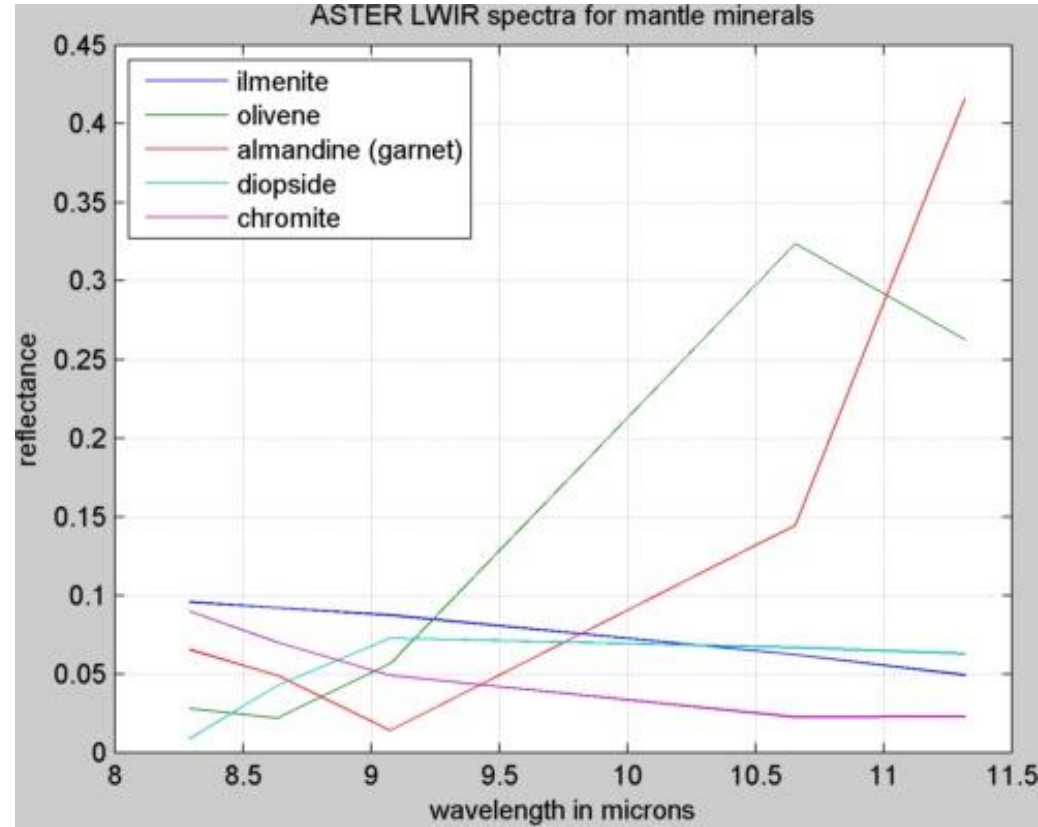
VNIR: 0.4-1.4 microns
SWIR: 1.4-3.0 microns
TIR: 8-11.5 microns

Size issues: 90m spatial does not give many pixels
5 spectral bands does not give many minerals

Kimberlites from mantle
along with chromite,
diopside, garnet, ilmenite
olivine

Very distinct LWIR spectra

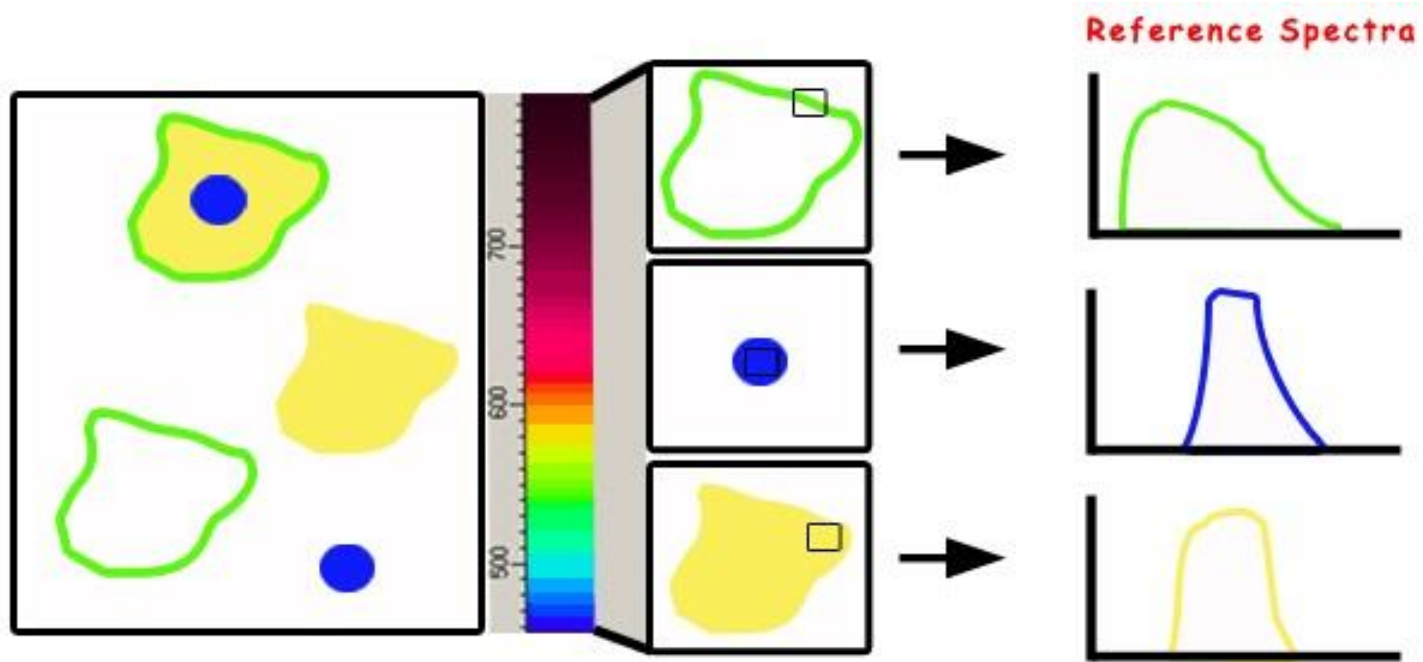
Components of spectral
mixtures



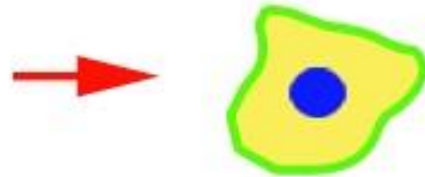
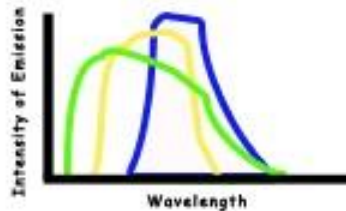
Spectral unmixing

Find endmembers
in scene
or external?

then unscramble
the egg



Linear Unmixing



Orapa diamond province
Aster VNIR image 7 May 2000

Legend

- Diamond Mine
- Feature 1
- Lethakane
- Orapa Police Station

8

40 km

N

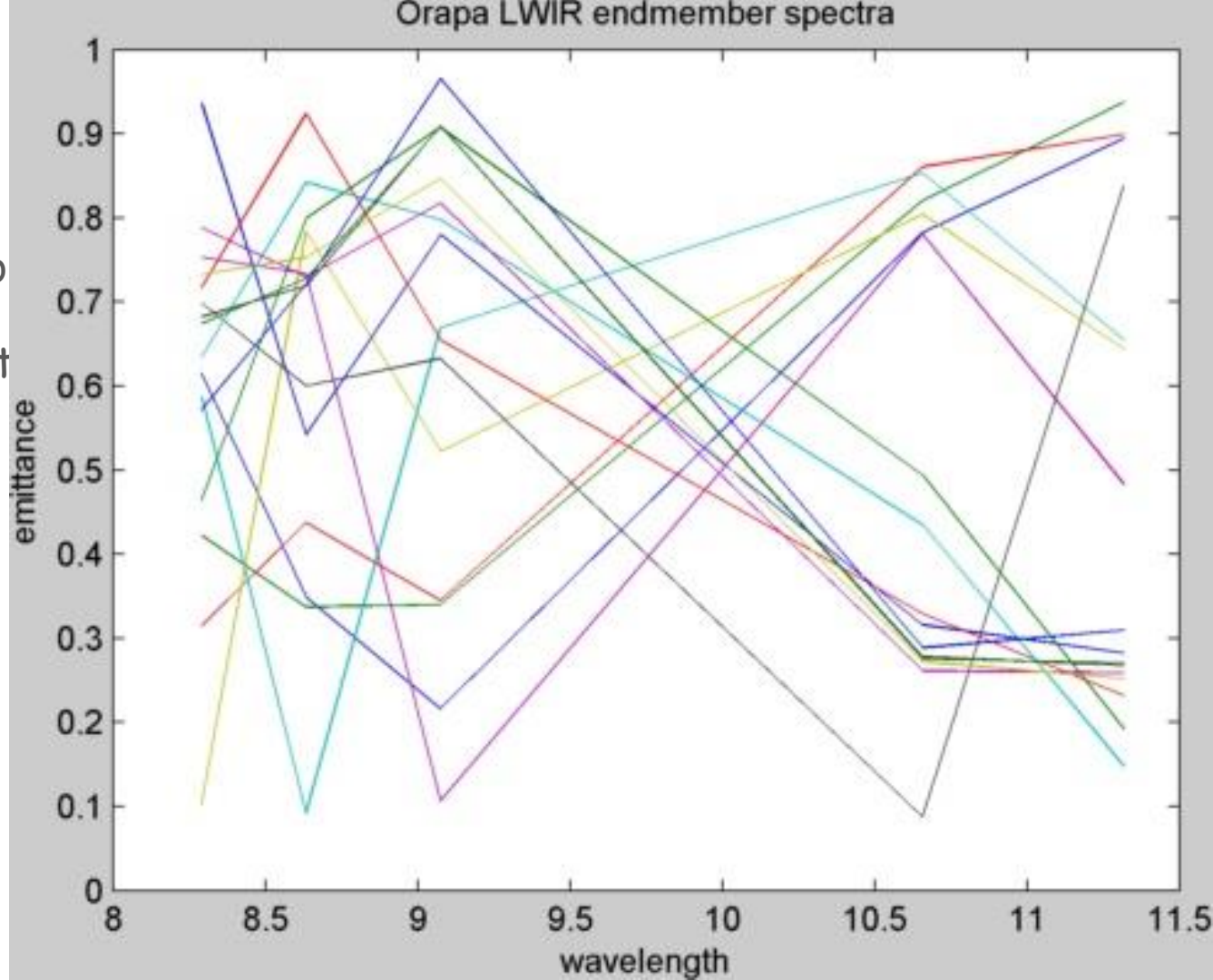
Processing flowchart

Atmospheric correction

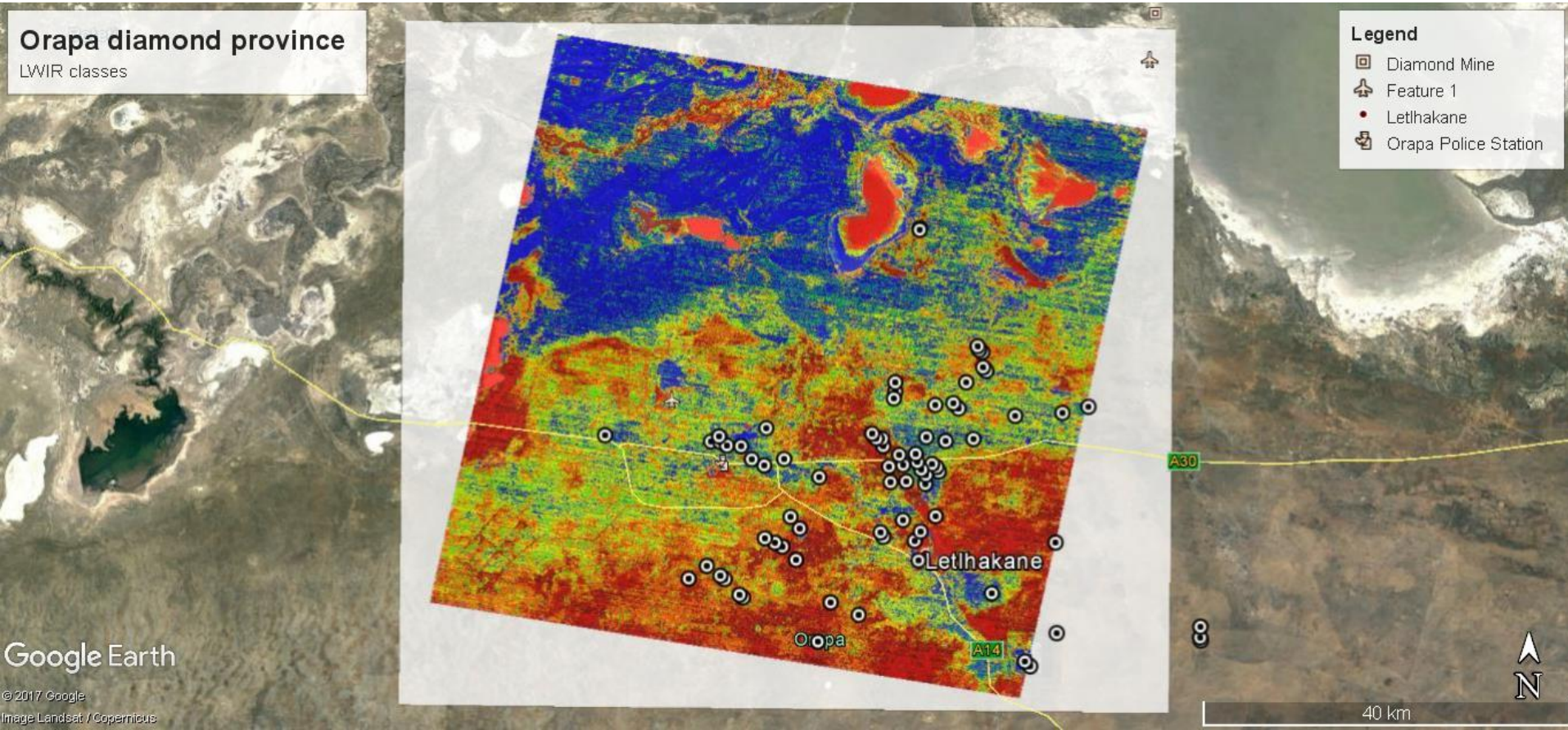
Temperature/emissivity
separation

Find 16 endmembers
to produce a spatially
sparse representation

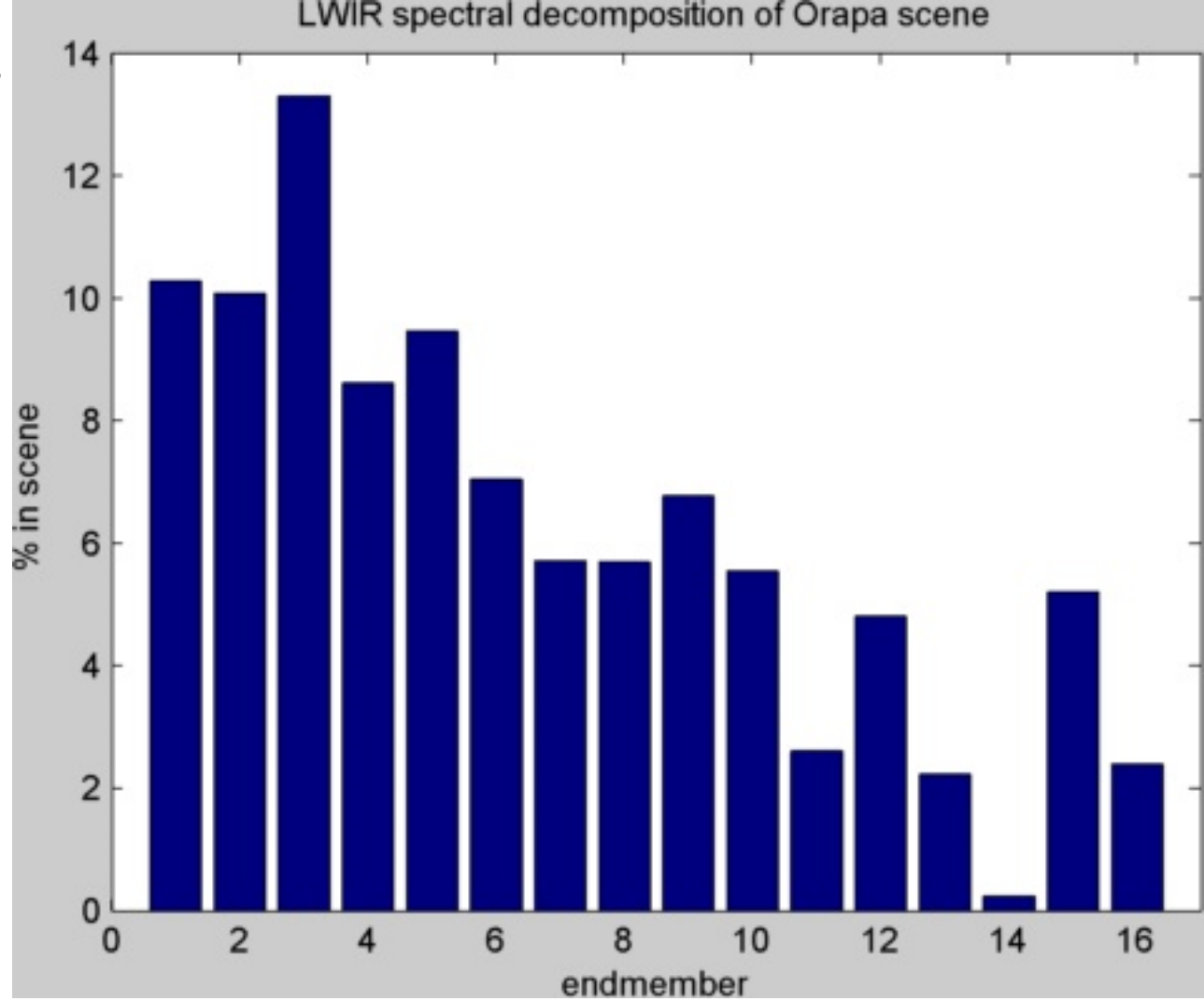
Project 5D LWIR
into a 16D space



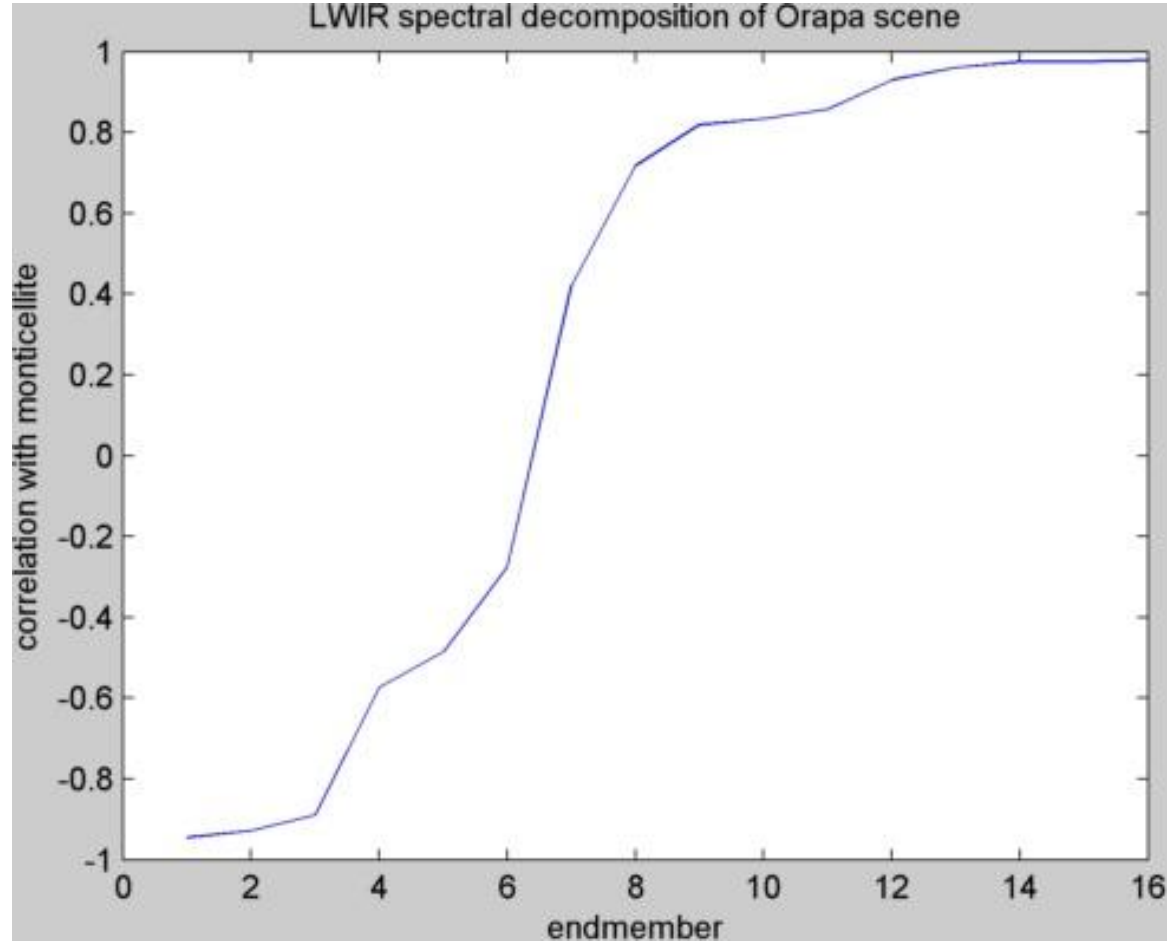
LWIR classes coloured 1=blue through to 16=red



Class abundances

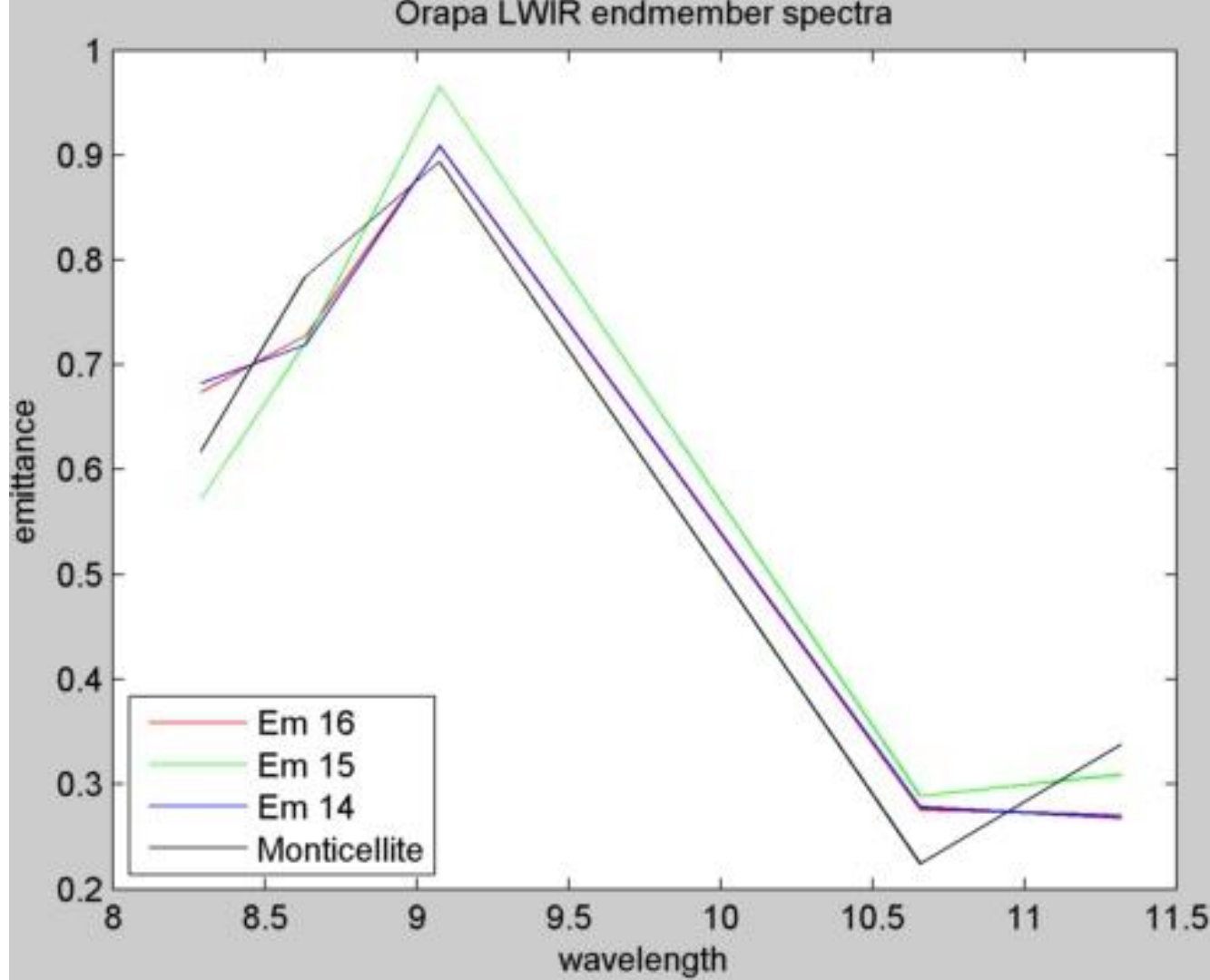


Sorted on their similarity to monticellite, a calcic-olivine

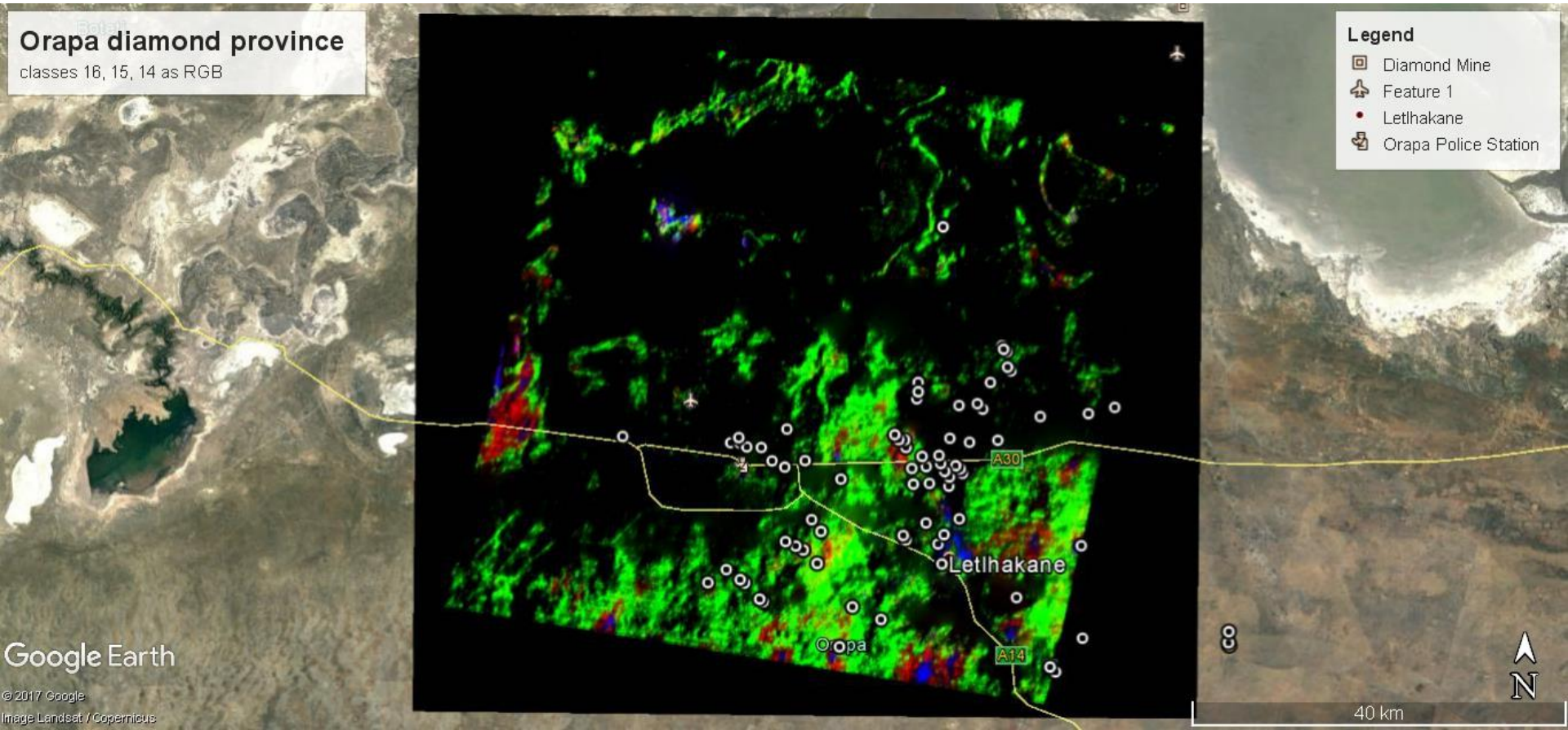


Top classes

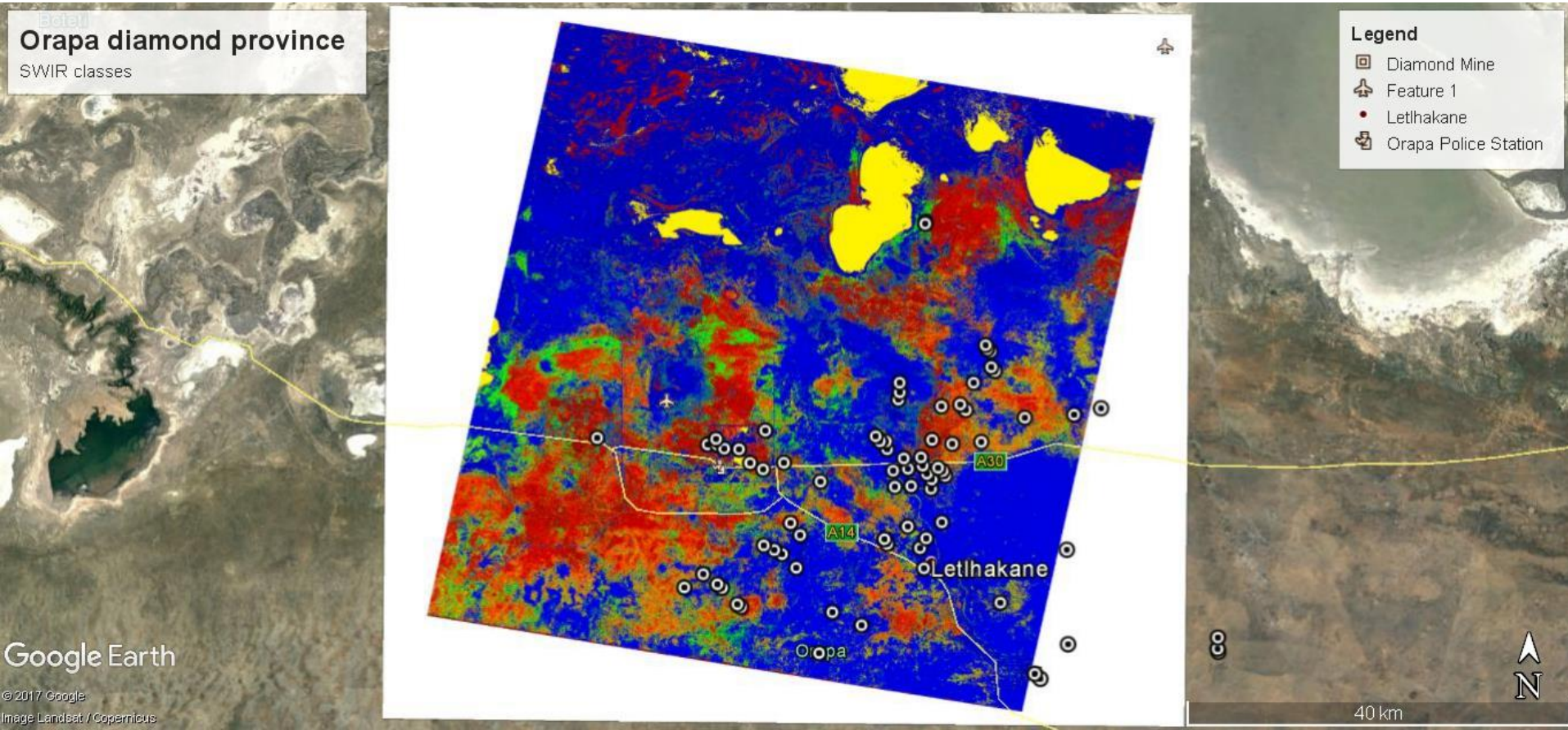
Emissivities
penetrate
vegetation and
windblown sand



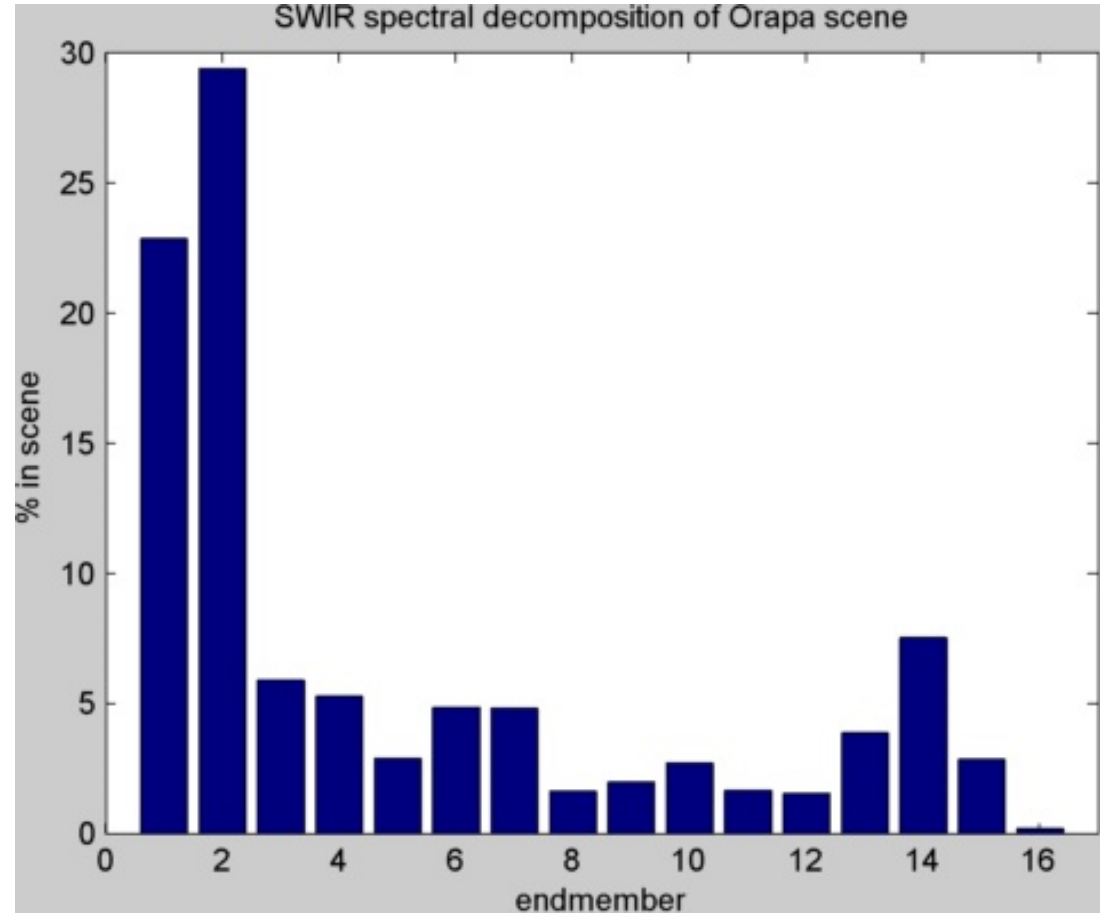
Top classes are spatially coherent



Can do the same thing with SWIR

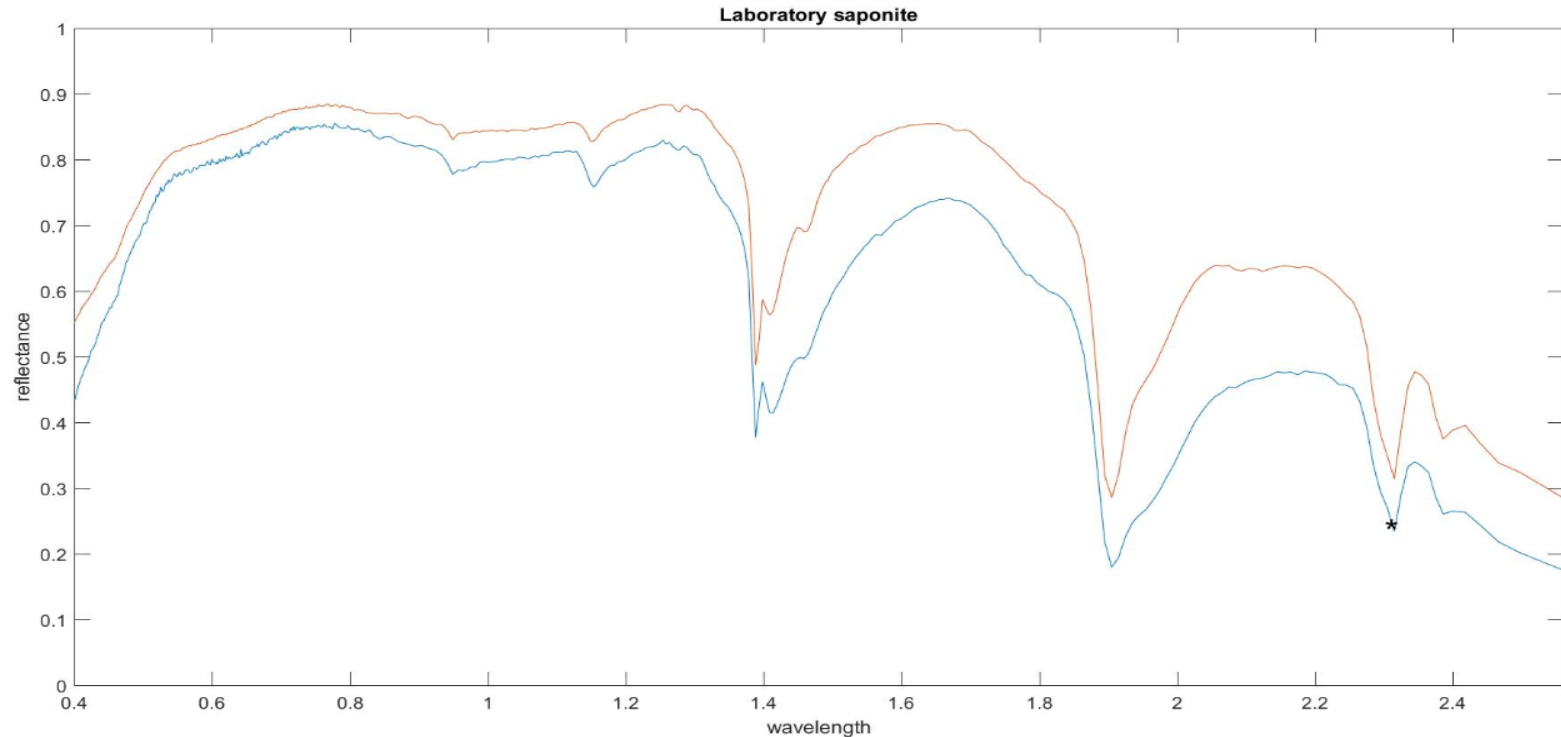


Class 16 (Saponite) maps the kimberlites



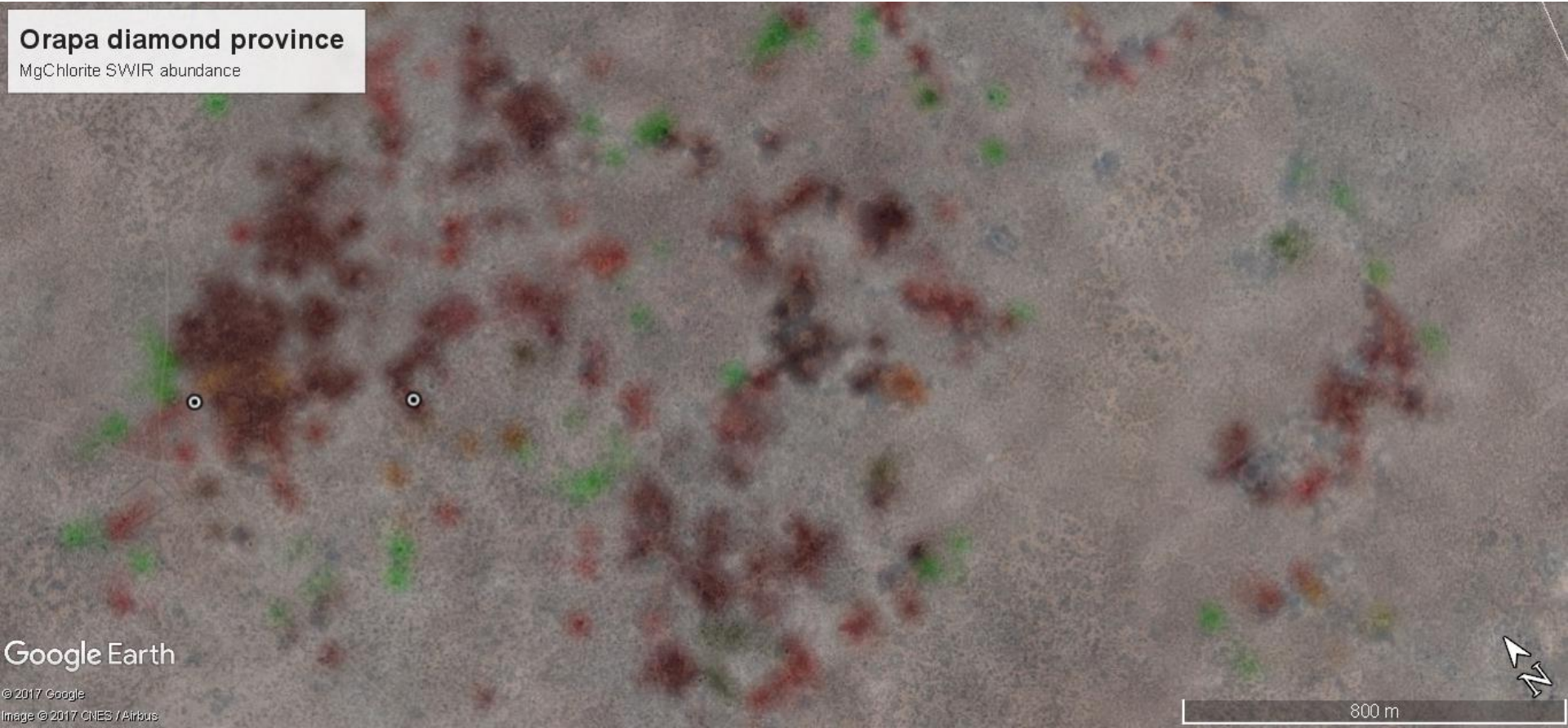
Saponite SWIR reflectance spectrum

focus on the 2.3 micron Mg feature, the reason De Beers pioneered hyperspectral kimberlite exploration

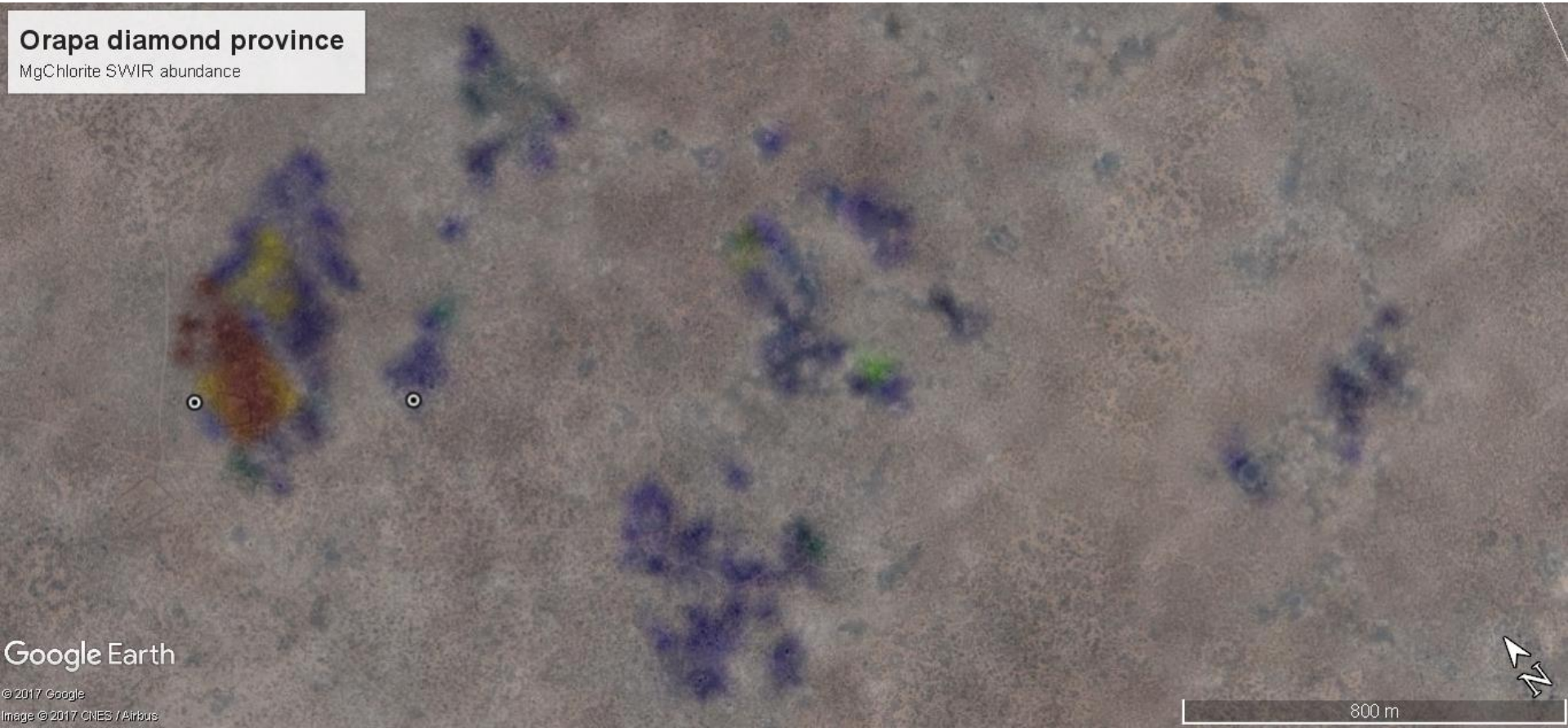


A scatter of saponite in the desert worth drilling?

○ are known kimberlites



Do a bit of Kriging

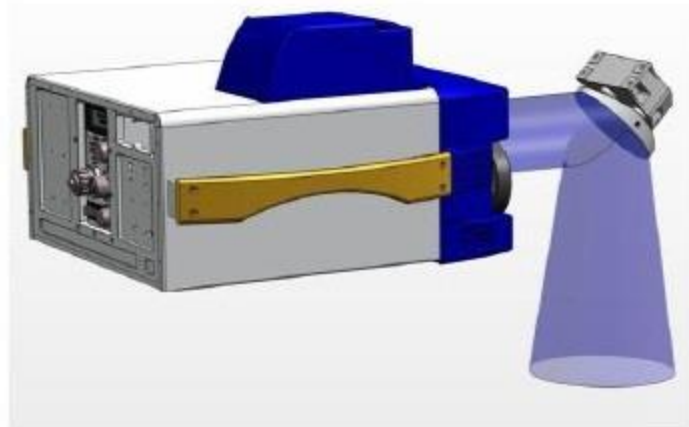


The next step: a Botswana group shoot

- Fourier transform infrared (FTIR) spectrometer
→ higher achievable SNR
- Michelson interferometer
- MCT focal plane array detector
→ adjustable acquisition area
- 2 internal calibration blackbodies
→ fast calibration
- Operability from -10°C to $+45^{\circ}\text{C}$
- Acceptable weight (30 kg)



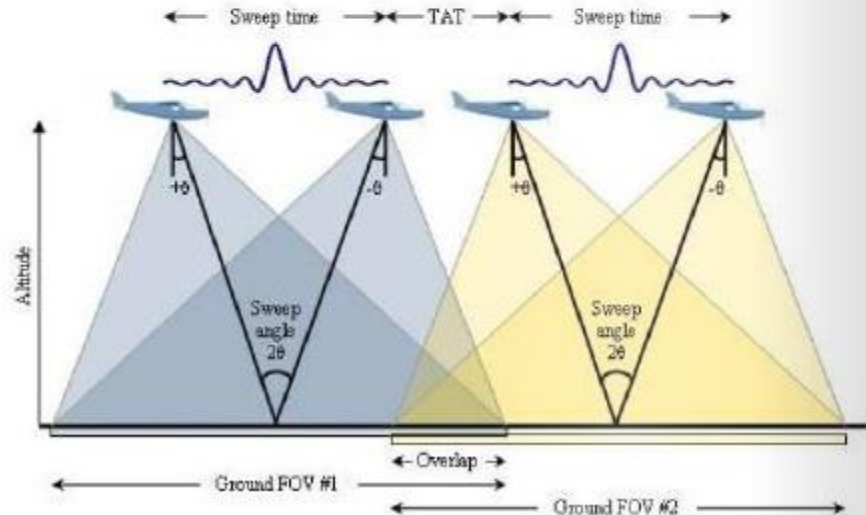
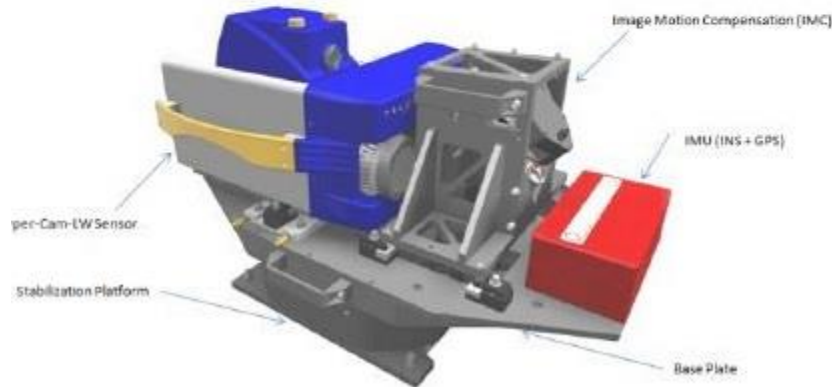
- Facilitates vertical measurements at ground level
- 45° tilted gold coated mirror that is located in the instrument's field of view
- 0.25x telescope
 - FOV at a sensor-target distance of 1.5 m is 672 x 538 mm
 - Resulting pixel size is 2.1 mm
- Airborne mode at 1500 m
 - FOV: 672 / 168 m
 - Pixel size: 2.1 / 0.53 m

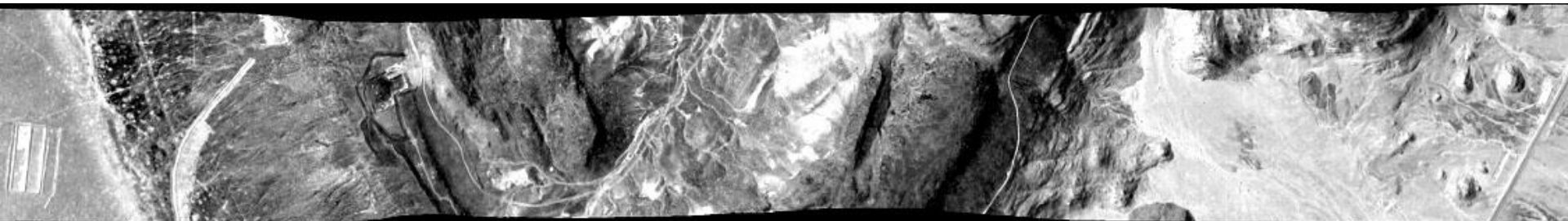
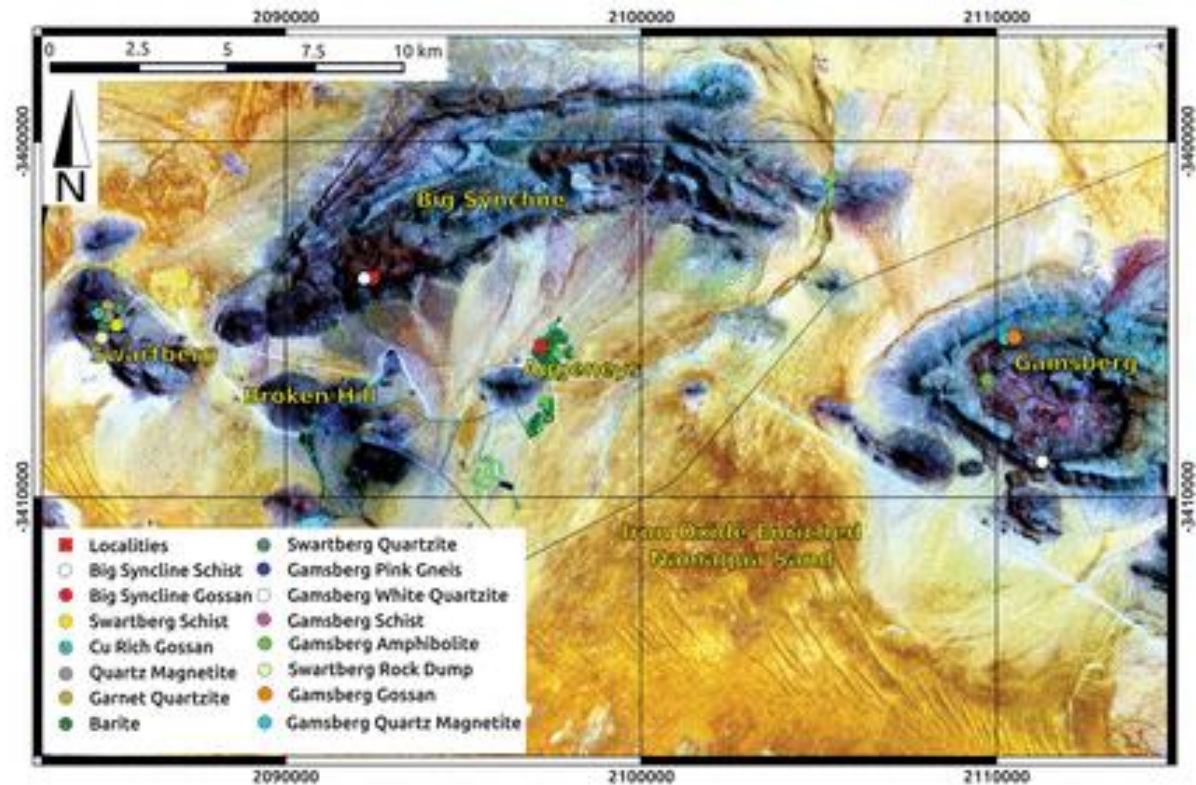


Hyper-Cam-LW specifications

Parameter	Unit	Hyper-Cam-LW
Spectral Range	μm	7.7 – 12
Spectral Resolution	cm^{-1}	0.25 to 150 (user adjustable)
Image Format	-	320 x 256 pixels
Field of View	Degrees	6.4 x 5.1 (nominal)
	Degrees	25.6 x 20.4 (0.25X telescope)
Typical NESR	$\text{nW/cm}^2\text{srcm}^{-1}$	< 20
Radiometric Accuracy	K	<1

- Stabilization platform: dampens the airplane vibrations and compensates the airplane yaw
- Image Motion Compensator (IMC) mirror: compensates the airplane pitch, roll and forward motion
- GPS/INS unit: enables ortho-rectification and geo-referencing

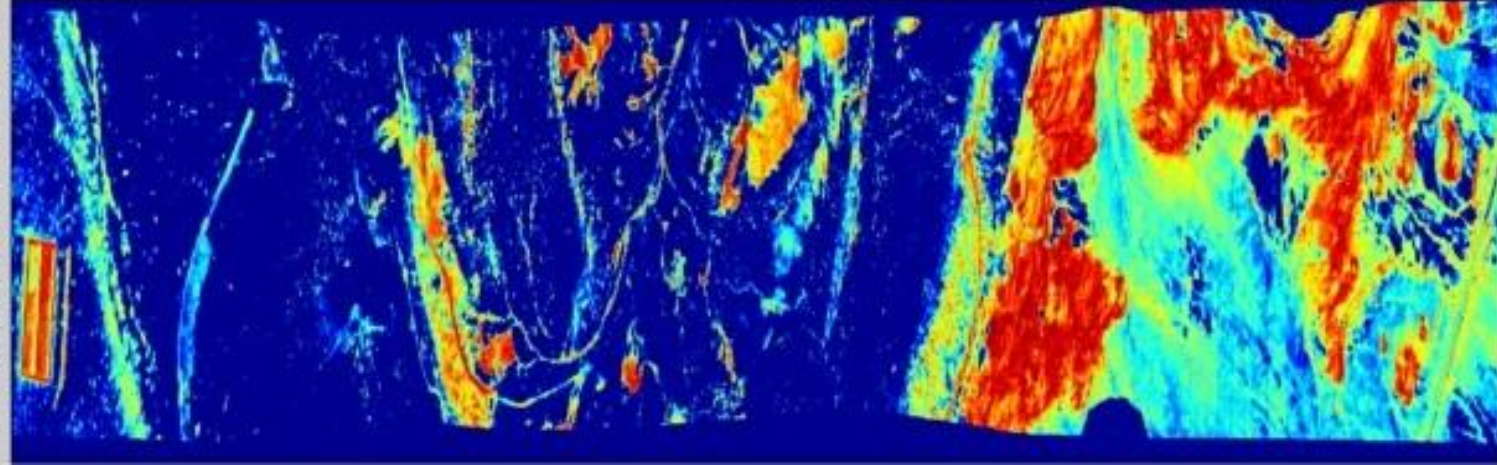




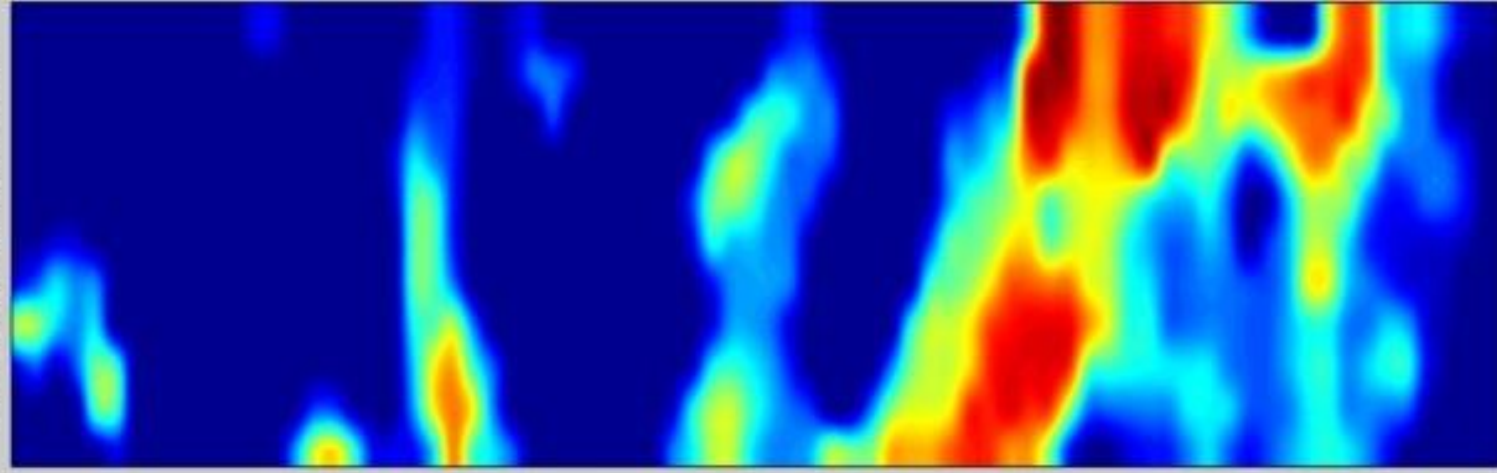
Spectral abundances



Telops thermal quartz 2.5m



Aster thermal quartz 90m



Conclusions

Mantle minerals associated with kimberlites can be rapidly and inexpensively mapped using ASTER LWIR imagery

The next step is airborne LWIR @2.5m spatial, 132 bands [7.6 11.4]

Remote geochemistry is a cost effective addition to the toolbox of the modern diamond explorationist

Good exploration targets in the southwest of the OKF

Zebediela kimberlite swarm worth investigating
as both Marsfontein & Klipspringer are LWIR anomalies